

Modeling the 40m

QND Workshop, Hannover

Dec 15, 2005

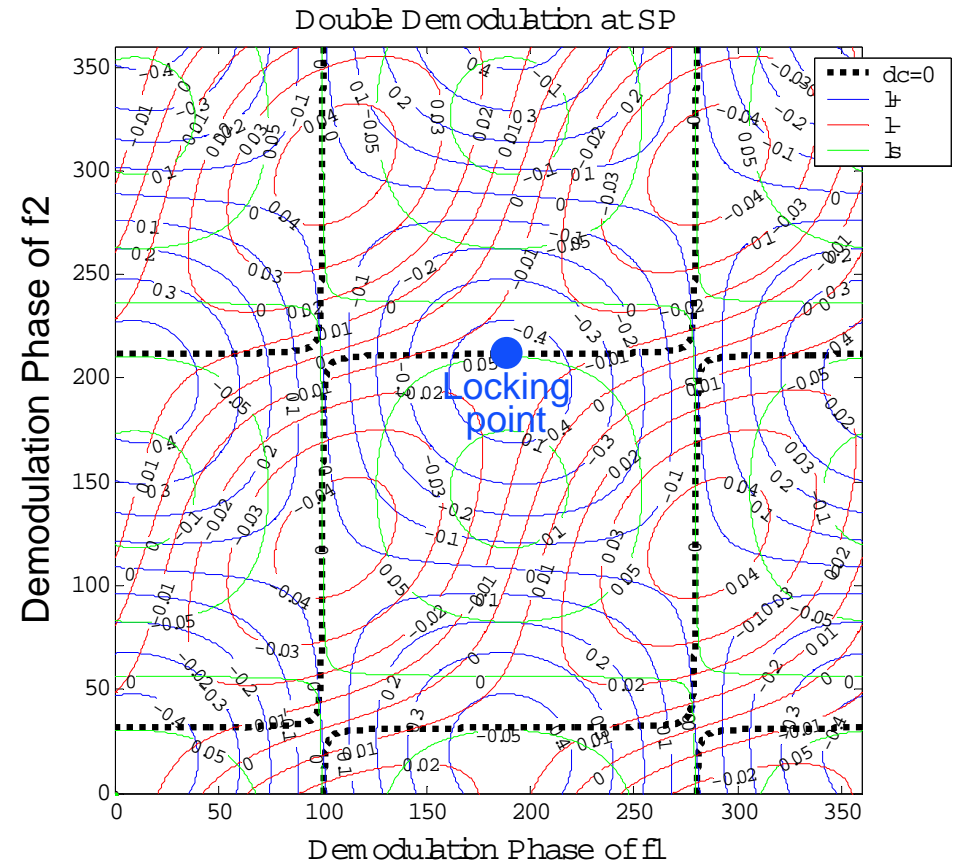
Robert Ward

Seiji Kawamura, Osamu Miyakawa, Hiro Yamamoto, Matthew
Evans, Monica Varvella

Modeling tools used at the 40m

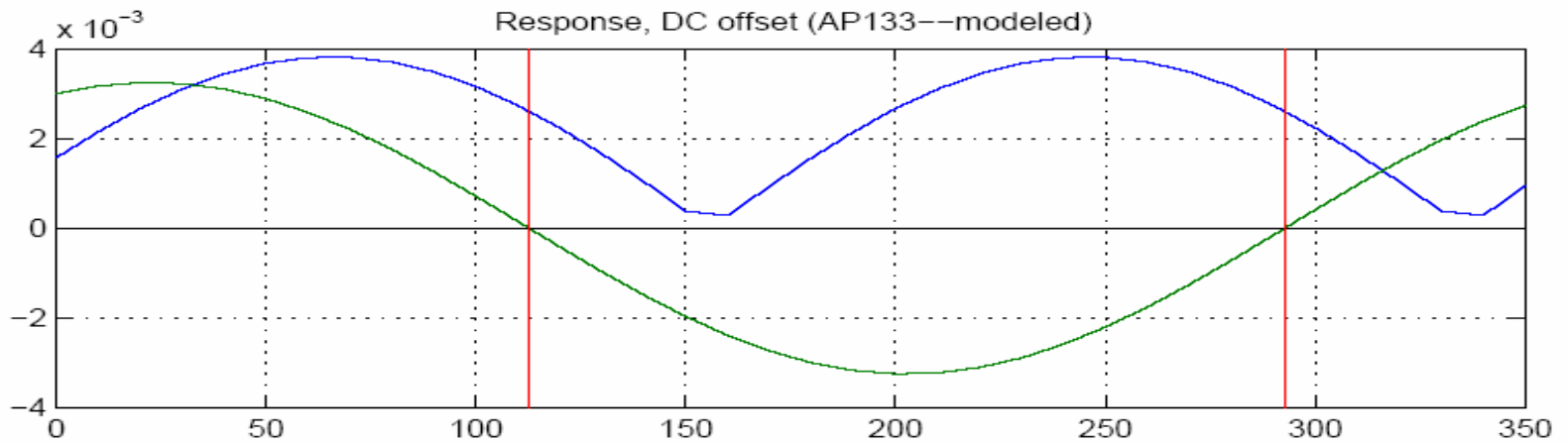
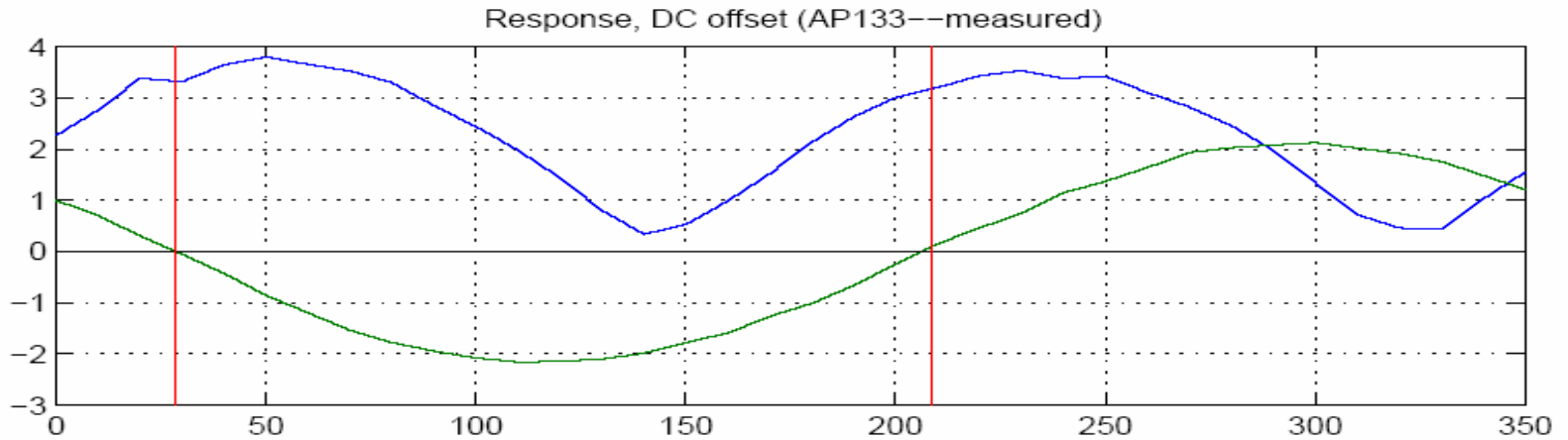
- Twiddle
 - » Frequency domain, analytical, no radiation pressure
- Finesse
 - » Frequency domain, no radiation pressure
- E2E
 - » Time domain, now with classical radiation pressure
- TCST (Thomas Corbitt Simulation Tool)
 - » Two-photon formalism, frequency-domain
- Optickle
 - » Frequency domain, Matlab, two-photon formalism

- Finesse used extensively at the 40m lab 2004-2005
- Seiji Kawamura modeled the DRMI very thoroughly using Finesse
- Very useful for investigating quirks of double & differential demodulation



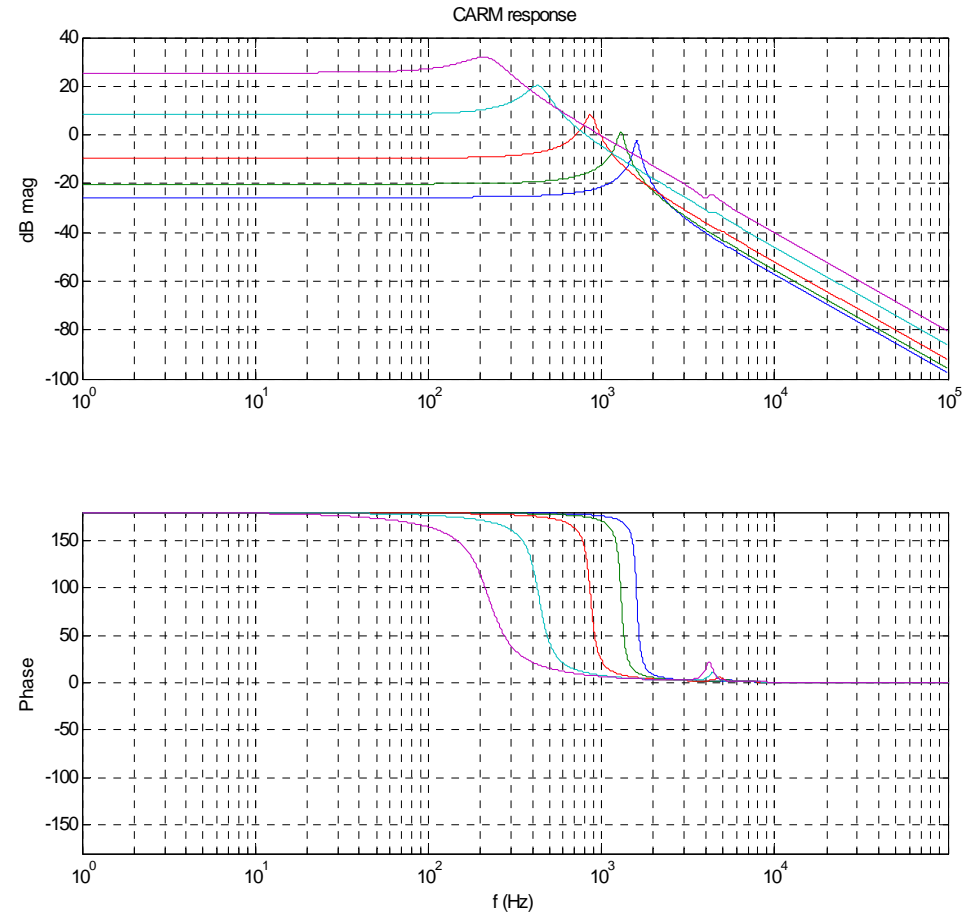


Differential Demodulation: offset vs gain



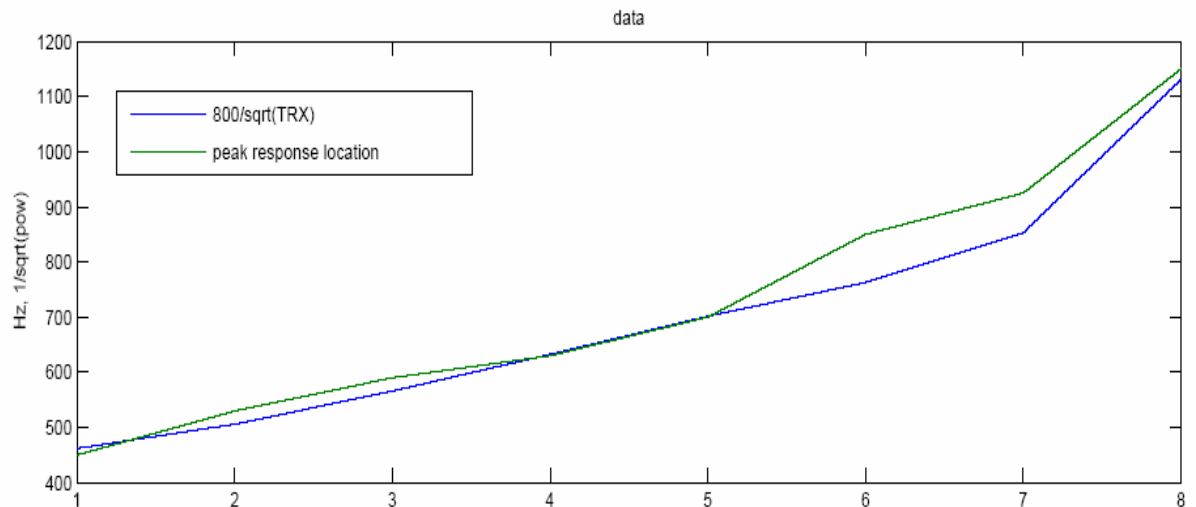
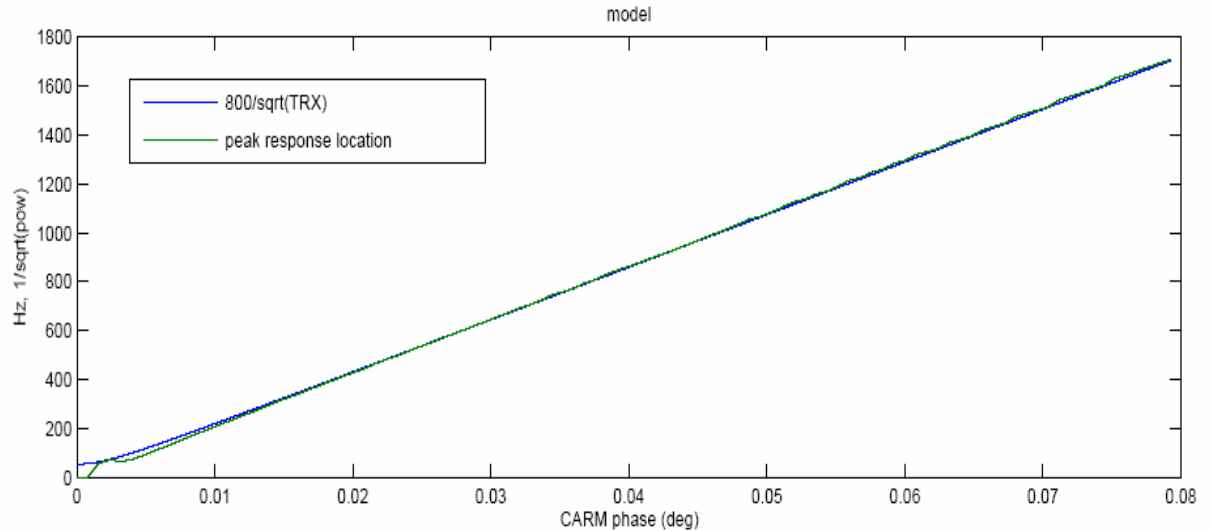
Finesse

- Also used to investigate the coupled cavity response in our offset CARM state, and design compensation.
- Unfortunately has no radiation pressure effects



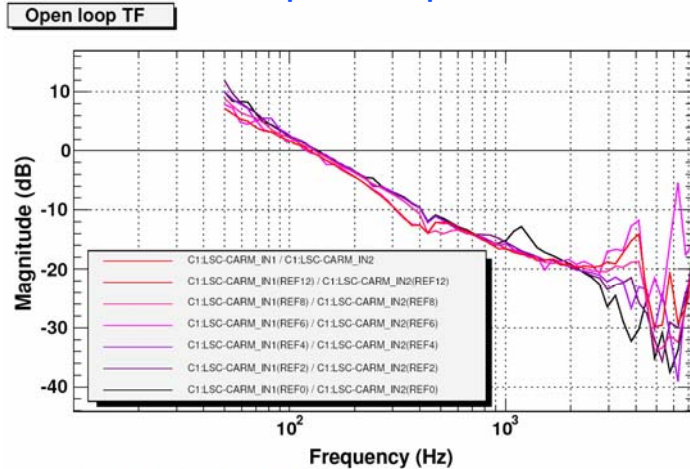
Straight from the ilog

- Location of peak RSE response (in CARM) as a function of offset, modeled in Finesse, and then measured.



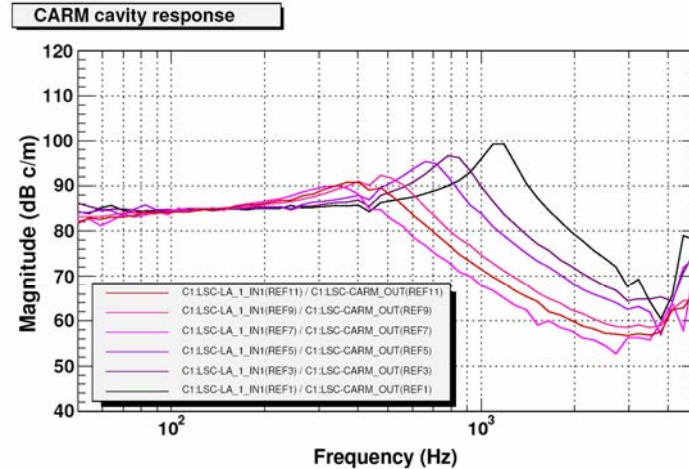
Dynamic compensation filter for CARM servo

Open loop TF of CARM



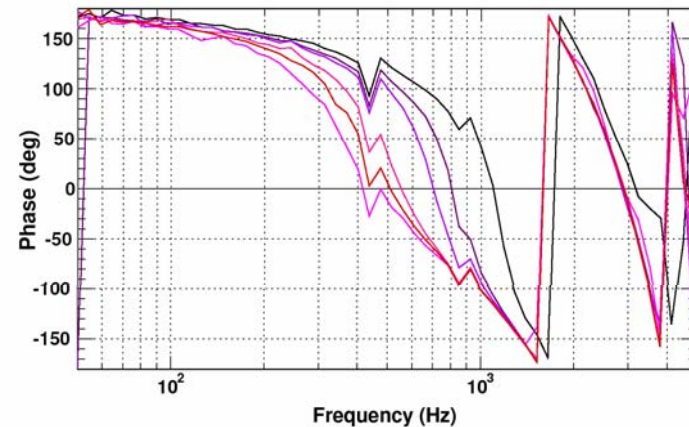
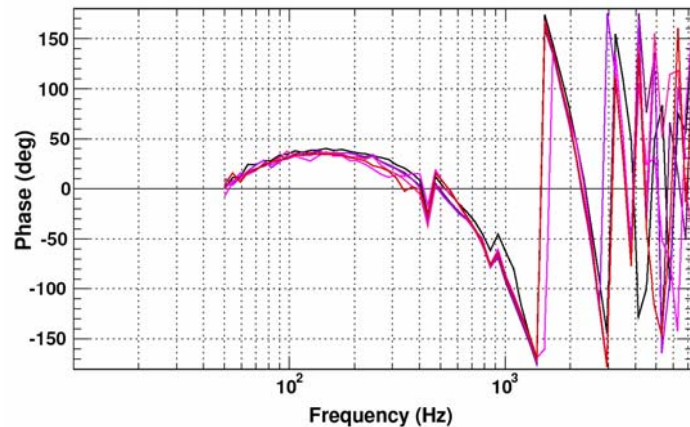
*T0=12/08/2005 12:10:00.040039 Avg=4

Optical gain of CARM



*T0=12/08/2005 12:08:00.040039 Avg=4

- Optical gain (normalized by transmitted power) shows moving peaks due to reducing CARM offset.
- We have a dynamic compensative filter having an nearly the same shape as optical gain except for upside down. Designed using FINESSE.



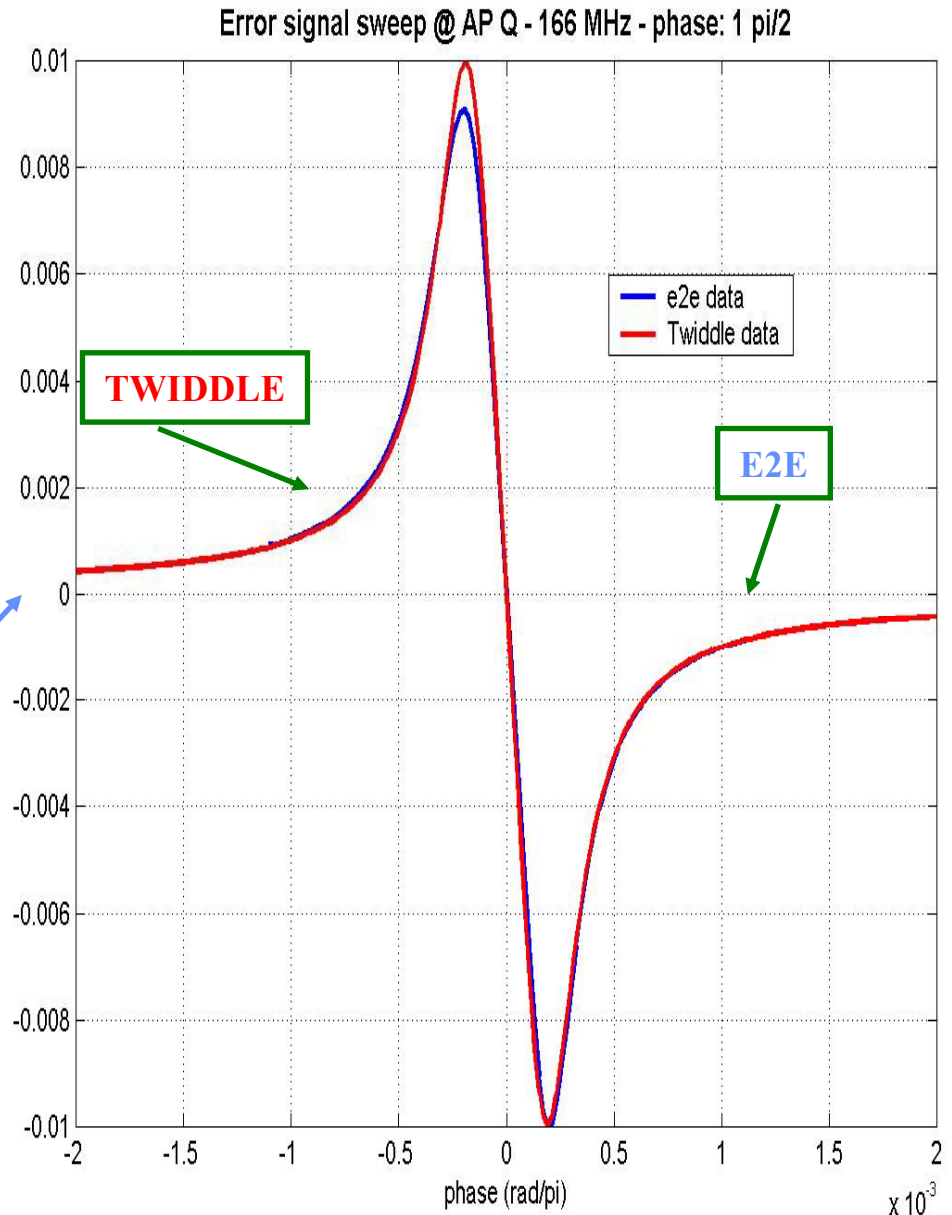
- Open loop transfer function has no phase delay in all CARM offset.

LIGO e2e SIMULATION:40m/AdvLIGO package

Error signal sweeps at 10^{-9} m/s
for the 40m IFO obtained in
E2E framework and compared
with TWIDDLE predictions

Example:
DARM @ AP 166 MHz
TWIDDLE and E2E
comparison

40m I

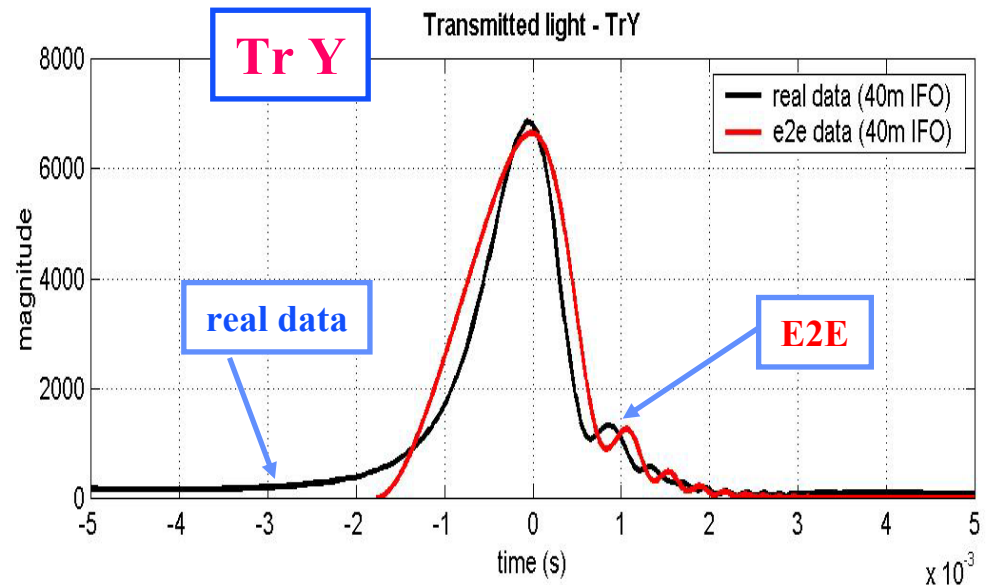
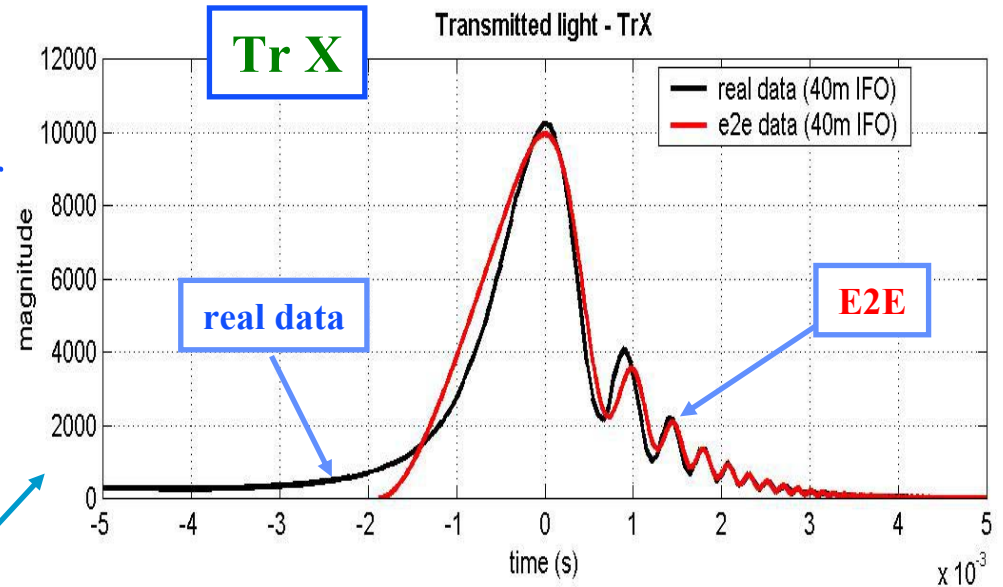


Real data have been used to estimate relative mirror velocity for both the arms:

$$V_{xarm} = (0.35 \pm 0.13) \mu\text{m/s}$$

$$V_{yarm} = (0.26 \pm 0.13) \mu\text{m/s}$$

Comparison between real data (black) and e2e simulated data (red) of the transmitted light for both the arms (full IFO): the mirror velocities used in E2E simulation are the values obtained fitting the real data



Comparison between real data ,
e2e simulated data and the
 theoretical prediction $V(t)$ of the
 SP error signal @ 166 MHz

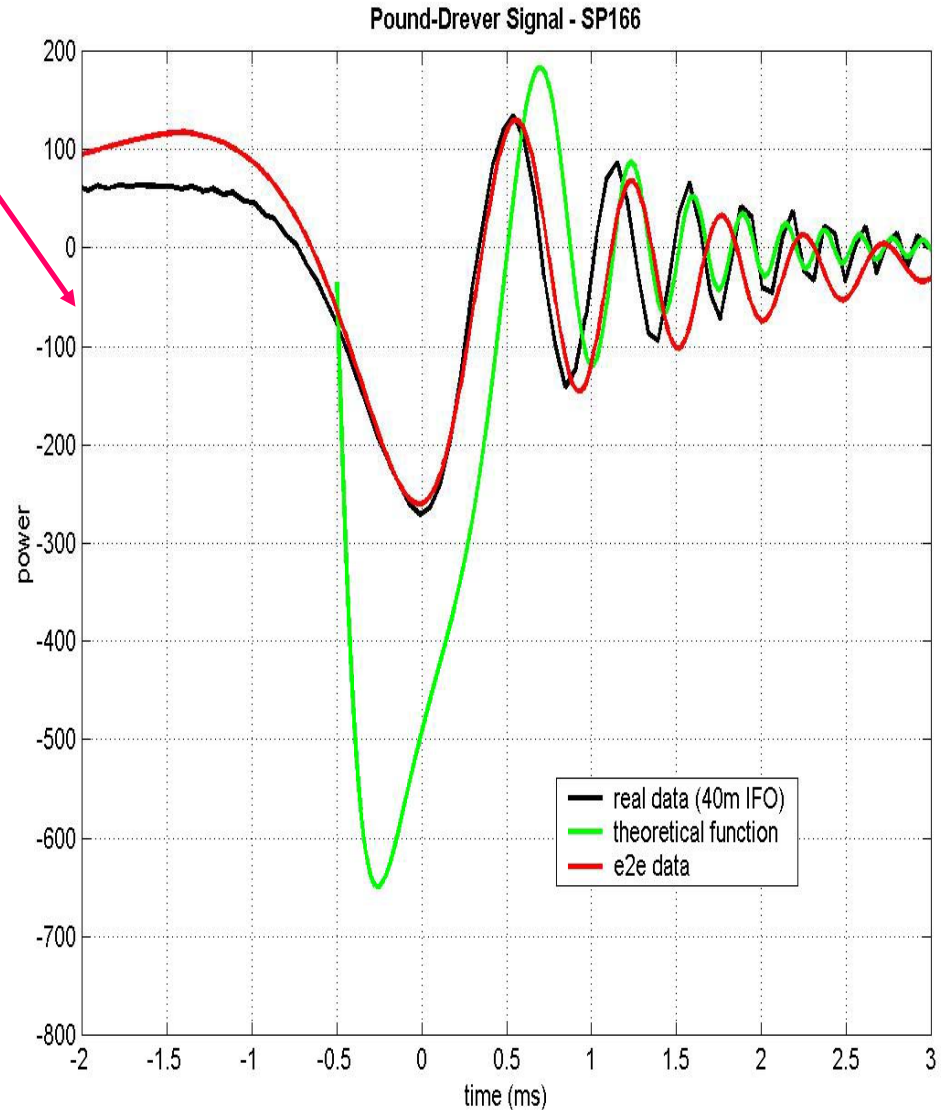
$$V(t) \sim \exp(t/\tau) \sin(a t^2)$$

with $a = (k v) / (2 T)$

The τ and the velocity v is the
 value obtained fitting real data

$$\tau = 0.7 \text{ ms}$$

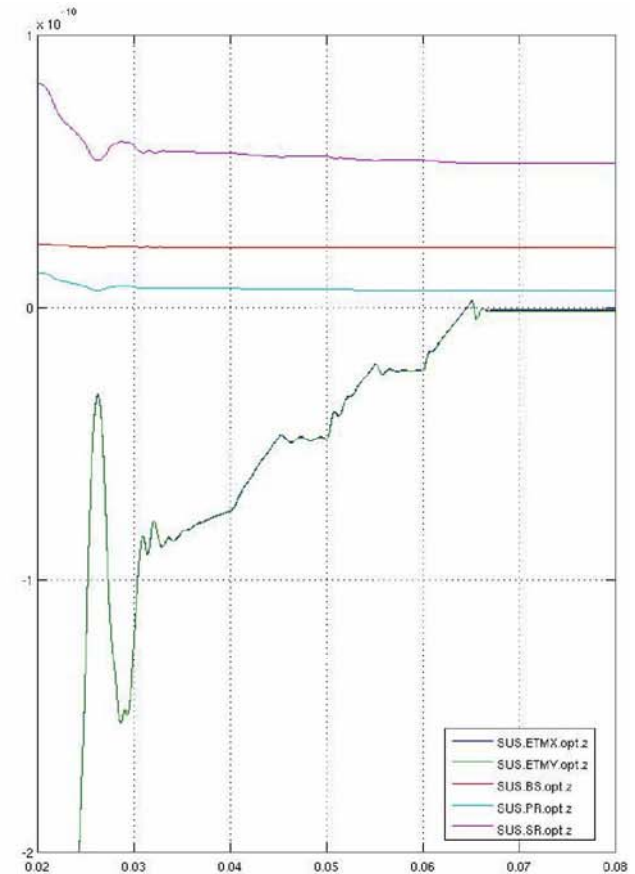
$$v = 0.26 \text{ } \mu\text{m/s}$$



E2E: 40m Lock Acquisition

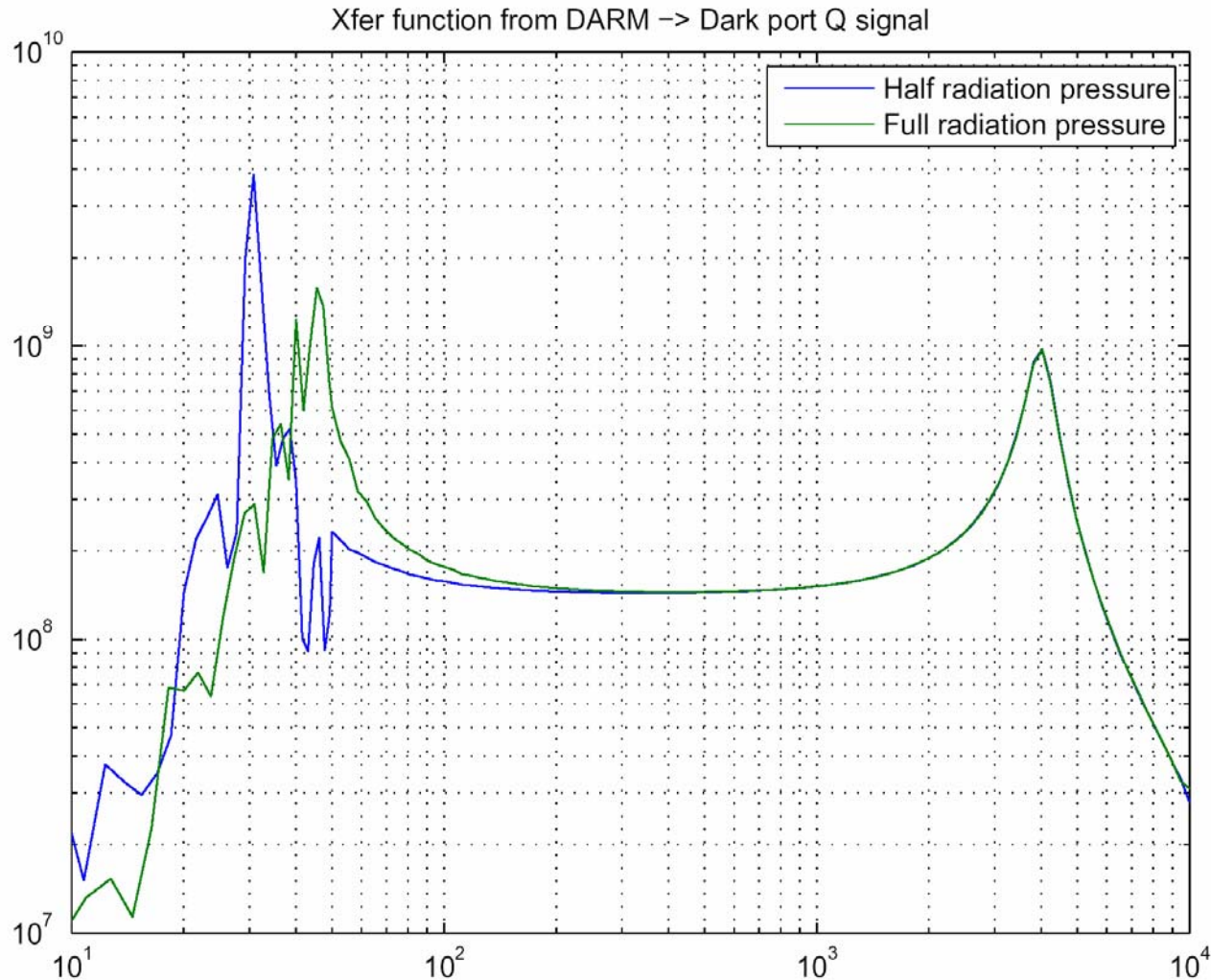
E2E simulation by Matthew Evans in June 2005

- ◆ CARM moved to RF signal
 - » Not yet done at 40m
 - » REFL port HF demod (a.k.a., SP166)
 - » Normalized by arm power
 - » Offset and gain matched at hand over
 - » Offset swept to zero slowly
 - » Coupled-Cavity pole compensation required
 - Pole (actually more complicated) moves down as resonance is approached
 - Compensation filter uses sum to make a “moving zero”
 - More detail may be required for 40m



■ Simulation indicates that controlled reduction of CARM offset should work.

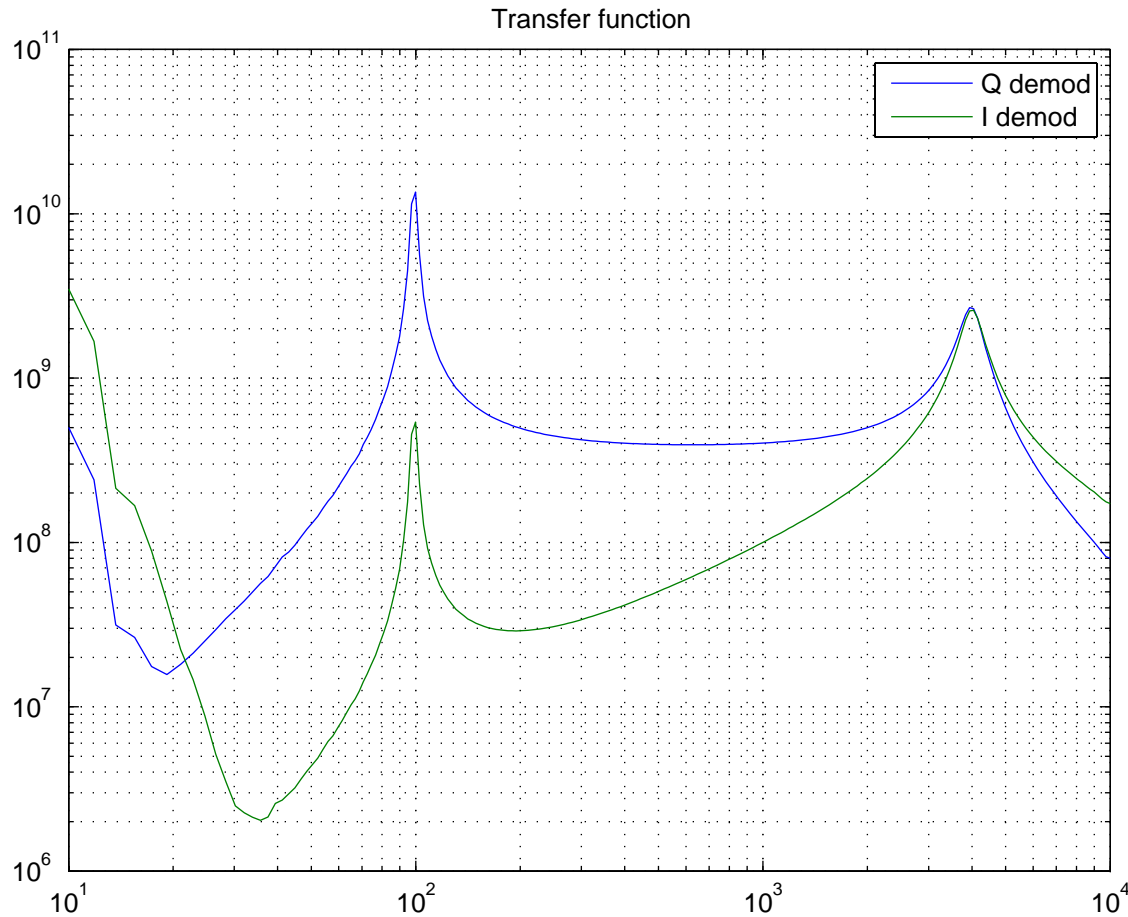
Optical spring in E2E



- Calculated by time domain simulation
- No length control
- Lock lasts ~0.7sec, so statistics at low frequency is not good.
- Simple length control required
- Calculation time ~5min using DRMI summation cavity

Hiro Yamamoto

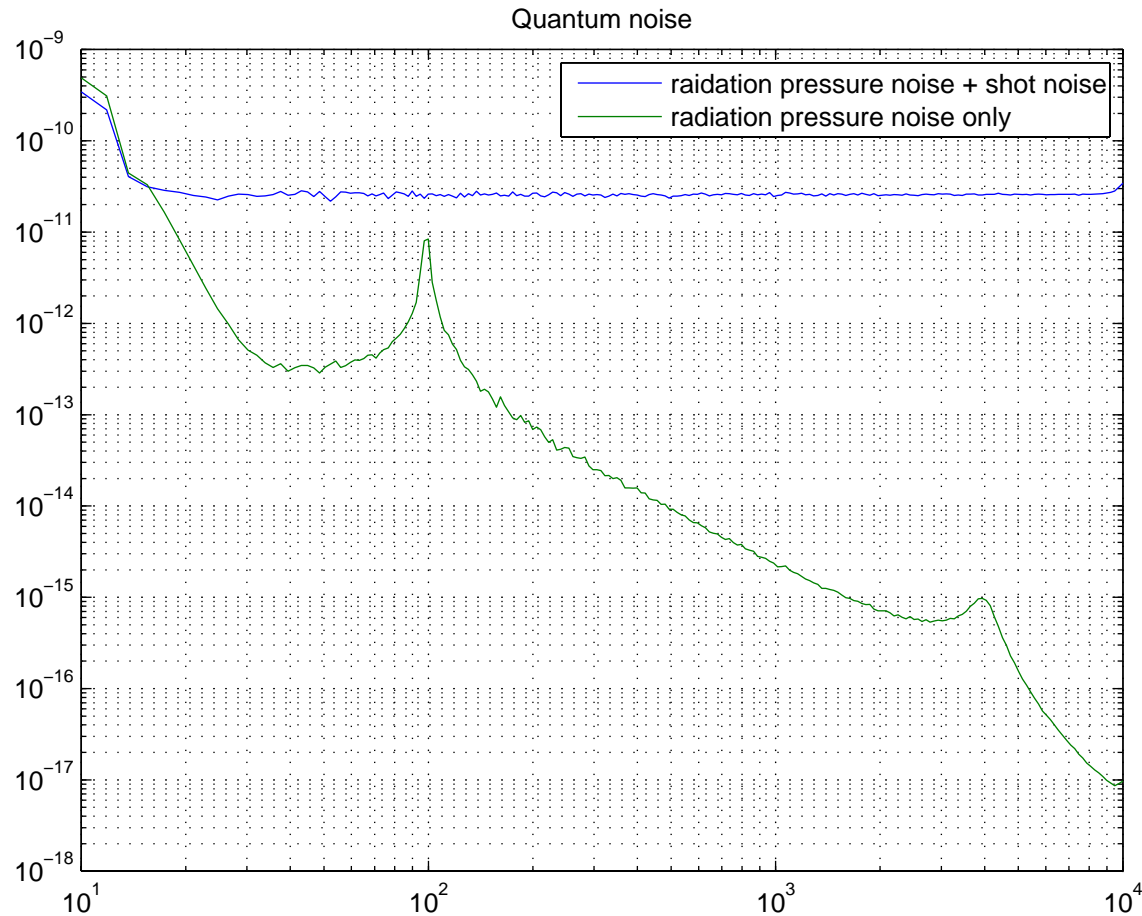
E2E DARM TF to I and Q



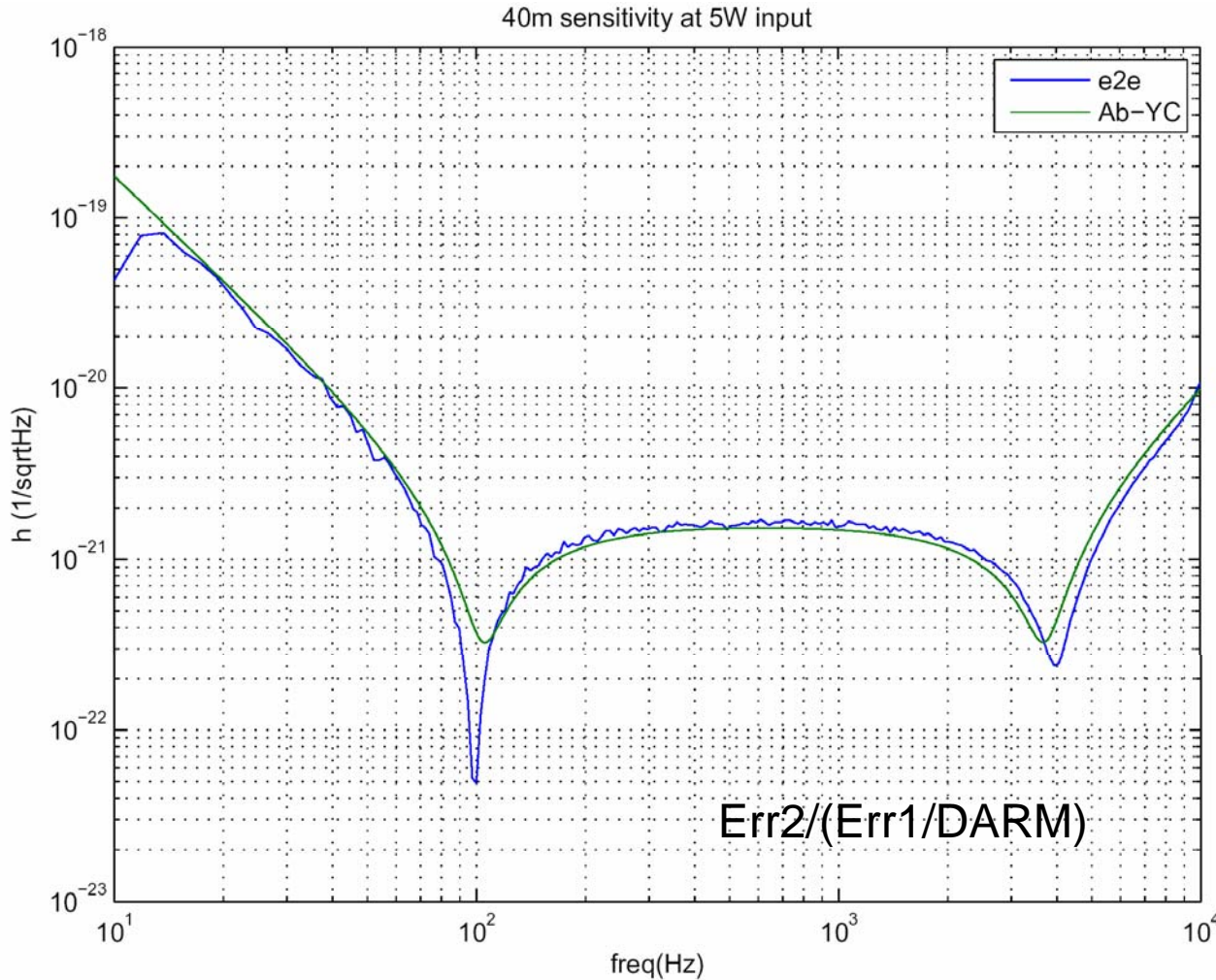
- 5W Input
- Arms controlled with POX, POY (no DARM)
- no MICH control

Hiro Yamamoto

E2E Optical Noise



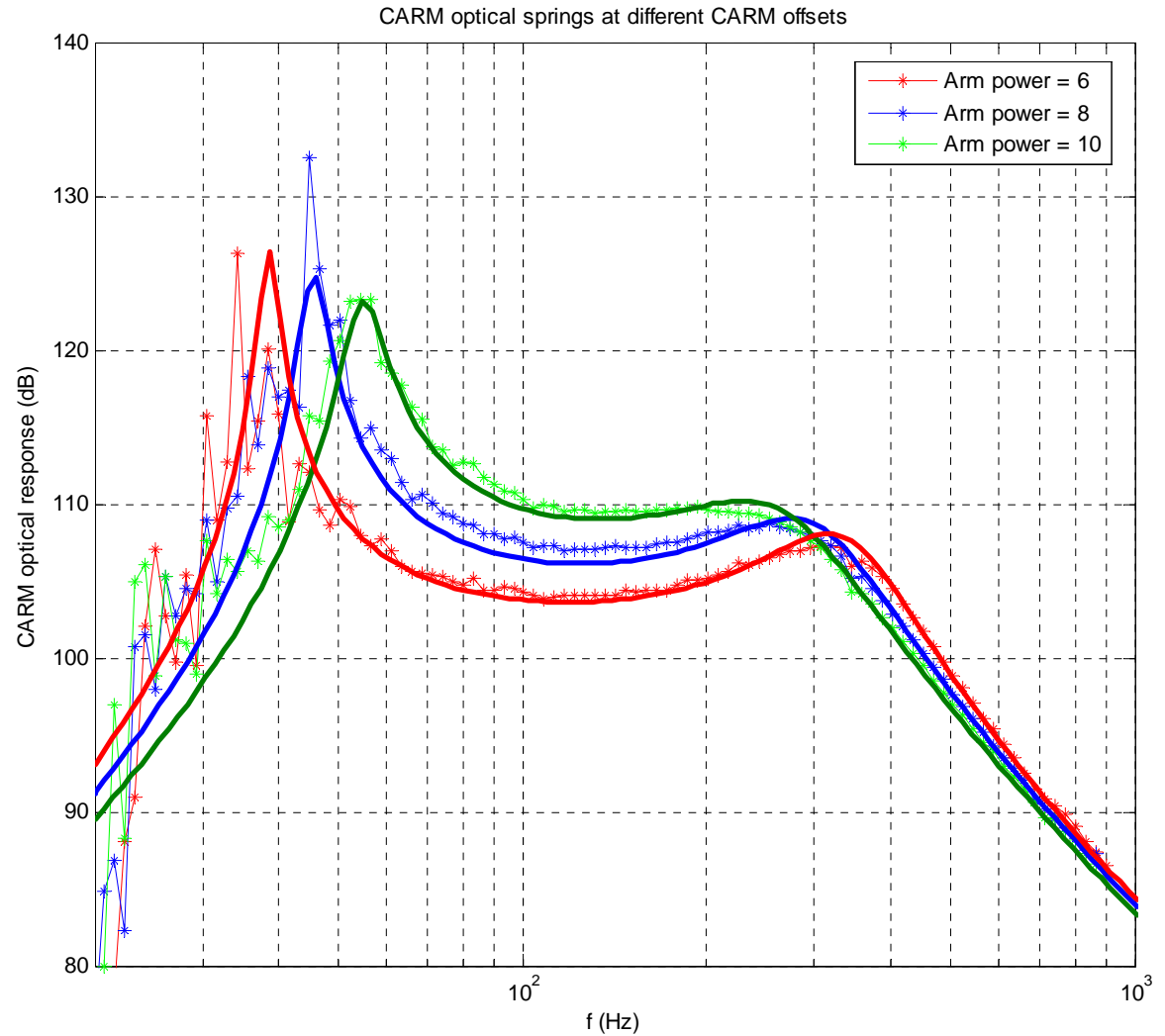
Optical noise of 40m in E2E



- Simple length control (UGF~100Hz)
- **Err2/(Err1/DARM)**
 DARM: DARM excitation on mirrors
 Err1: error signal with DARM excitation
 Err2: error signal with optical noise
- How much further does E2E need to go?
 2-photon?
- input vacuum?
- Quantum control?
- Or just classical physics + shot noise + radiation pressure noise ?

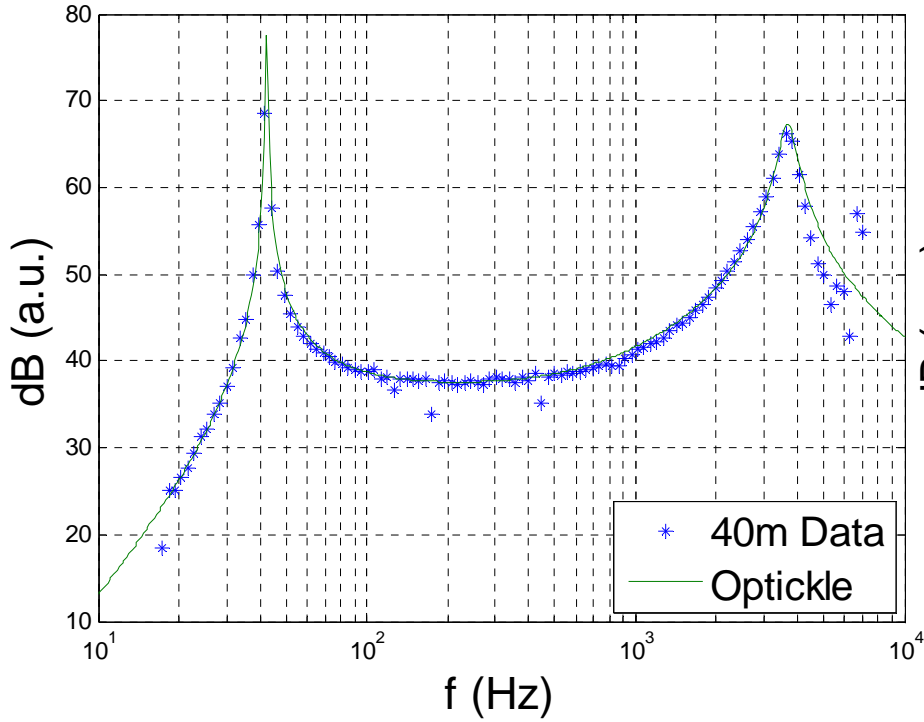
TCST

- Solid lines are from TCST
- Stars are 40m data
- Max Arm Power is ~80
- Also saw CARM anti-springs, but don't have that data

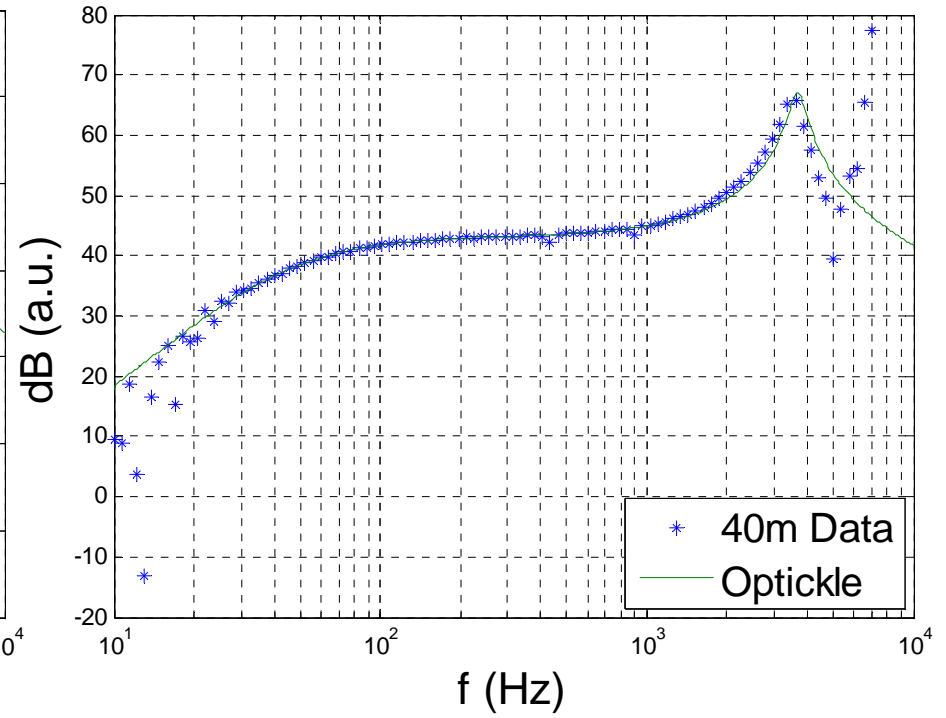


Optickle

DARM Response



DARM Response



CARM optical springs, with no offsets (TCST)

